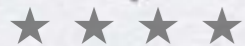


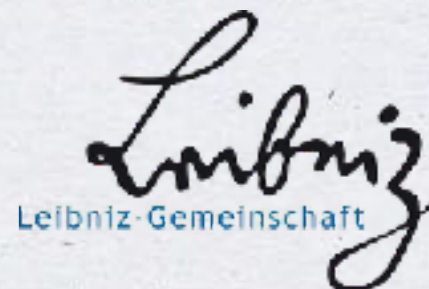


# TOWARDS A SCIENCE DATA CENTER FOR EST

Morten Franz



European Data Provider & Training Event  
Heidelberg, June 27<sup>th</sup> & 28<sup>th</sup> 2018





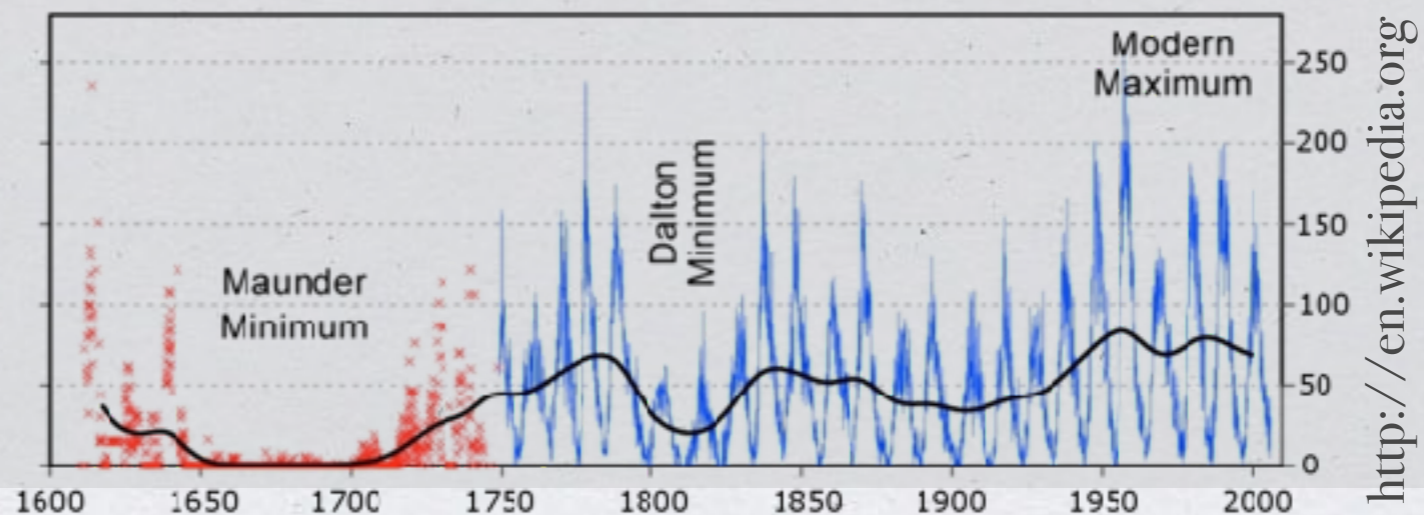
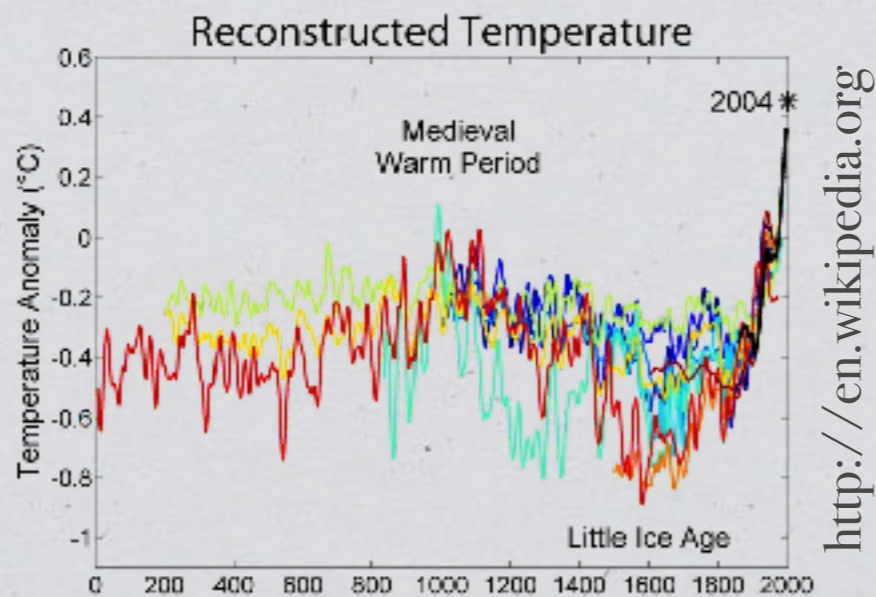
# Contents

- \* Solar physics and ground-based solar observation
- \* Upcoming solar telescopes → EST
- \* A flavor of solar data
- \* Summary



# Solar Physics

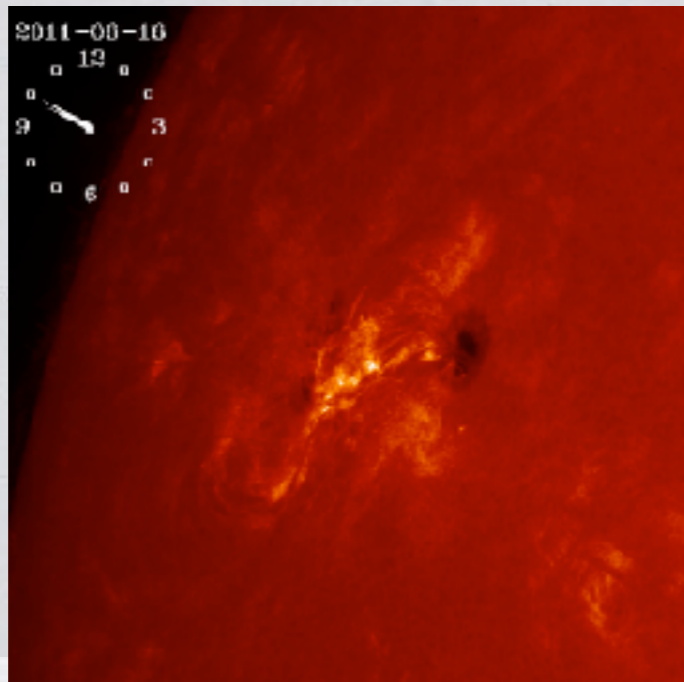
- \* The Sun has a significant impact on Earth's daily life
  - Solar cycle, sunspot number and solar activity



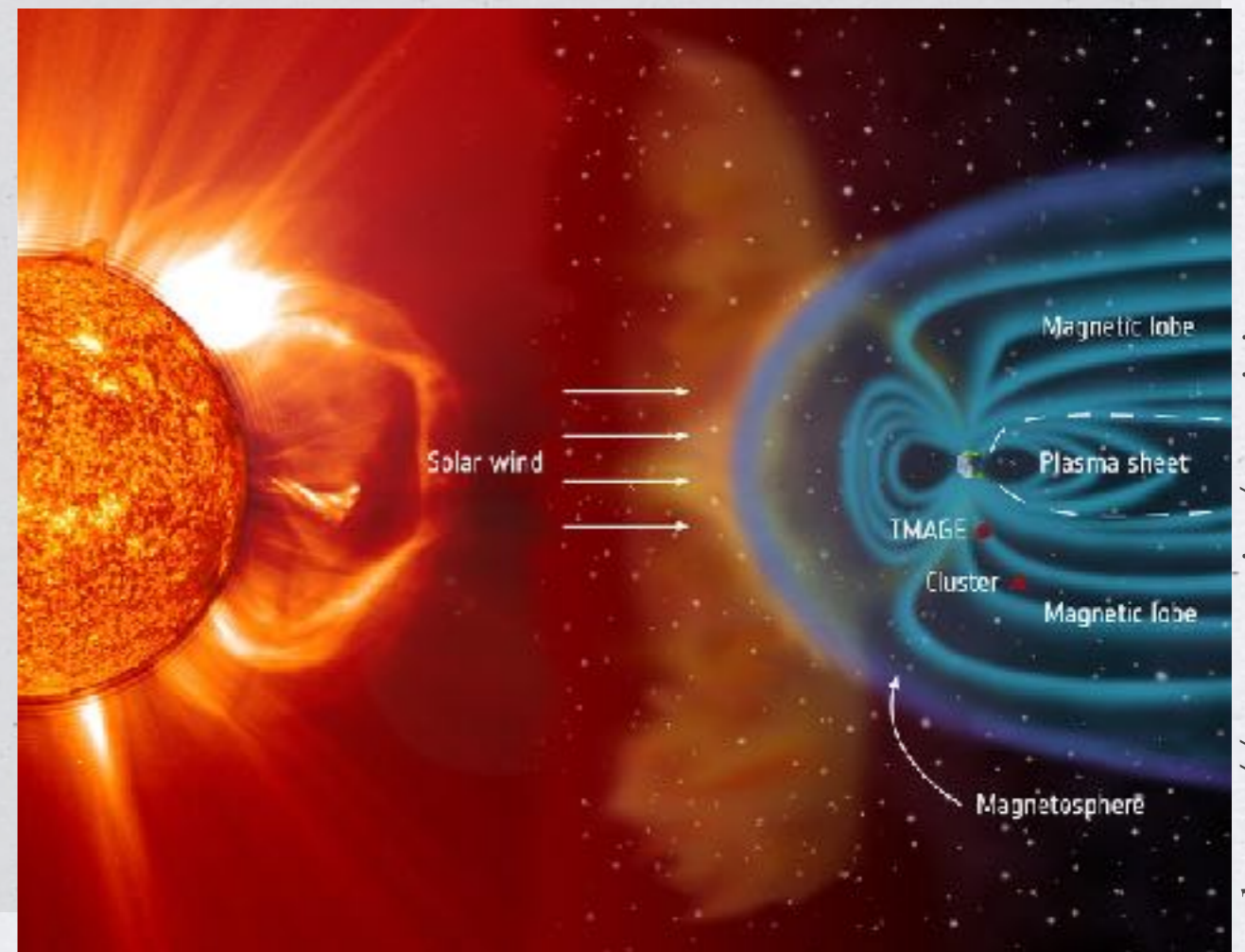


# Solar Physics

- \* The Sun has a significant impact on Earth's daily life
  - ▶ Solar cycle, sunspot number and solar activity
- \* Space weather has an impact on Earth and its magnetosphere
  - ▶ Consequences for satellites, i.e. GPS
  - ▶ Rerouting of commercial flights
  - ▶ Power grid failure

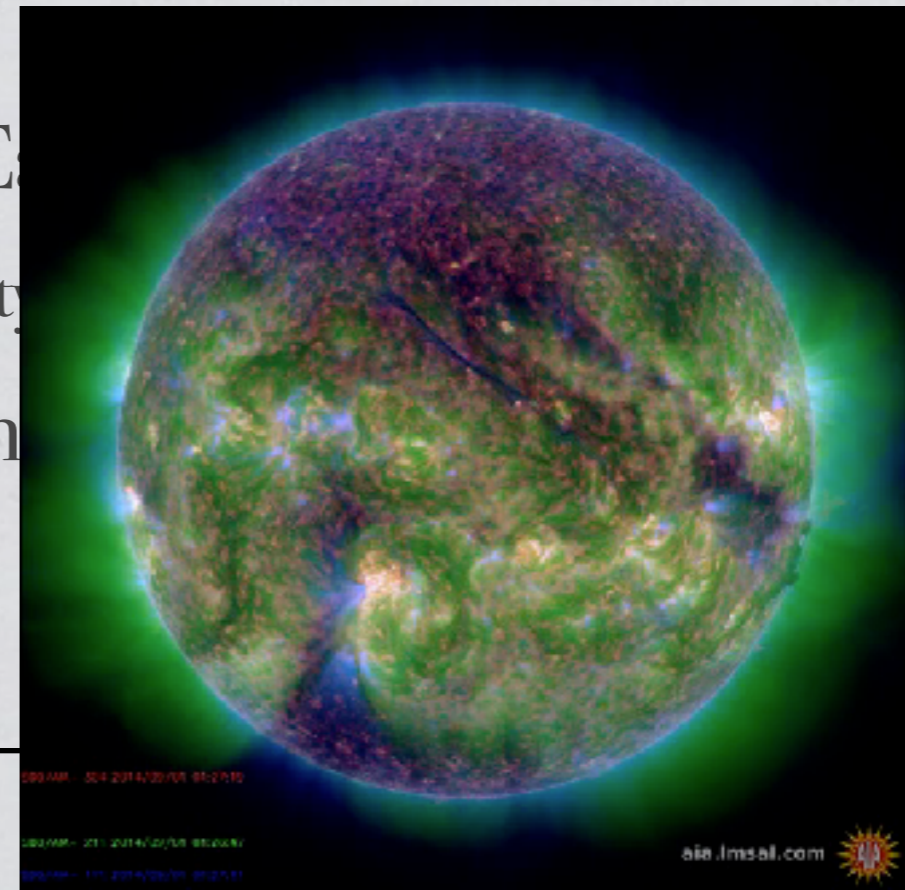
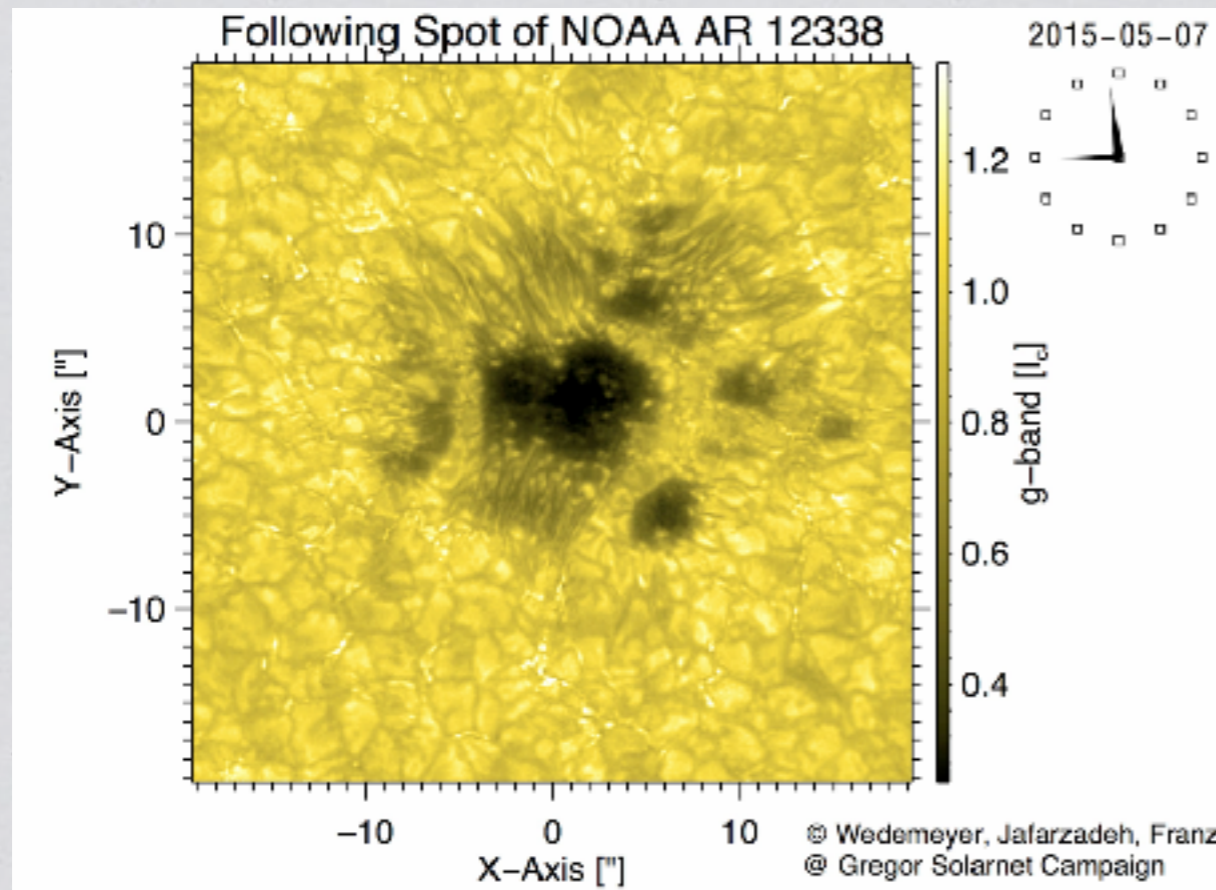


AIA 304 Å & AIA 4500 Å @ SDO





# Solar Physics



- \* Solar physics provides valuable input to the more general field of astrophysics
  - ▶ Sun is 'Rosetta' stone: Only star for which surface structures can be resolved in detail
  - ▶ Unique laboratory for magnetohydrodynamics
  - ▶ Abundance of elements



# Space-based vs. ground-based

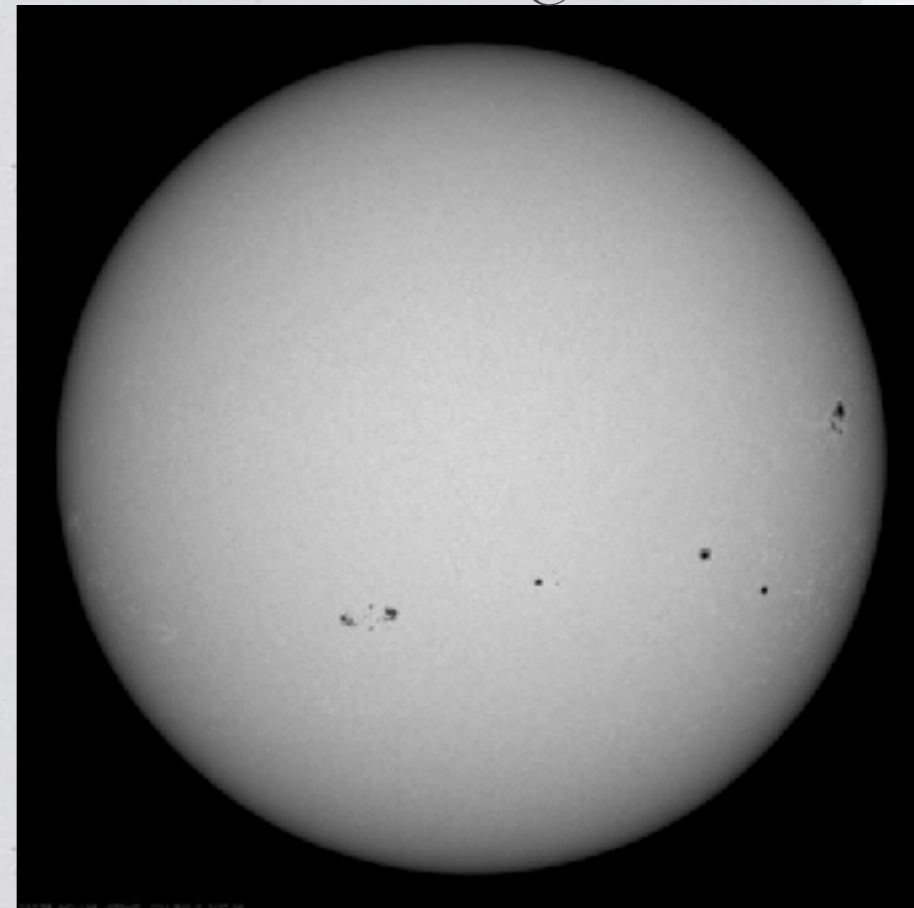
## \* Advantages of space-based observation

- ☑ Uninterrupted
- ☑ Uninfluenced by Earth's atmosphere
- ☑ Fixed observation modes
- ☑ Homogeneity of observational data

## \* Disadvantages of space-based observation

- ⚡ Inflexibility of observation
- ⚡ Inaccessibility of hardware
- ⚡ Limited in size
- ⚡ Extensive costs

HMI 4500 Å @ SDO





# Space-based vs. ground-based

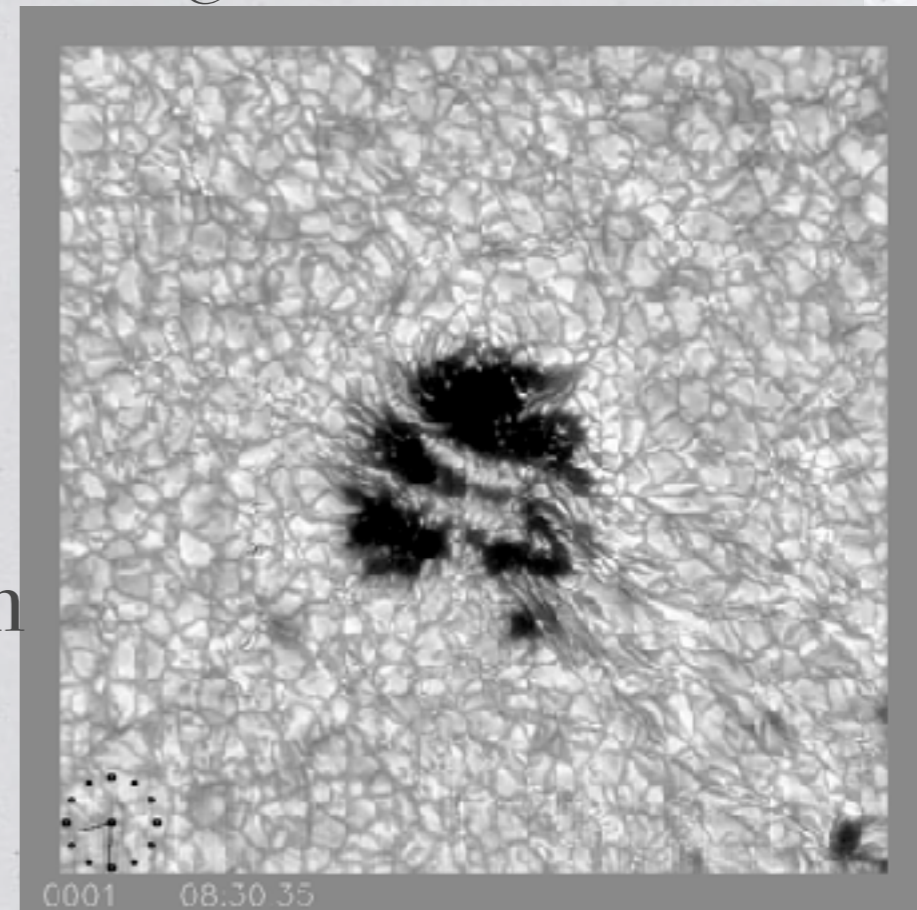
## \* Advantages of ground-based observation

- ✓ Flexibility of observation
- ✓ Hardware development during operational phase
- ✓ Cheap and easy maintenance
- ✓ Large facilities realizable at affordable cost

## \* Disadvantages of ground-based observation

- ✗ Interrupted by day-night cycle
- ✗ Influenced by Earth's atmosphere
- ✗ Currently PI based and run without standardized observing modes
- ✗ Heterogeneity of data

GFPI@VTT Schlichenmaier et al. 2010





# Upcoming Solar Telescopes

DKIST 2020



GST

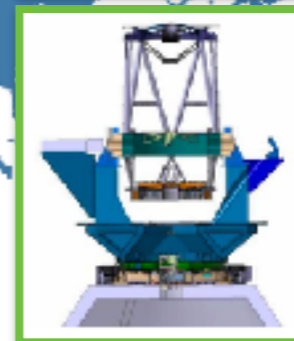
EST 2027



GREGOR, VTT,  
SST, Themis



Sayan Solar  
Telescope  
Coronagraph

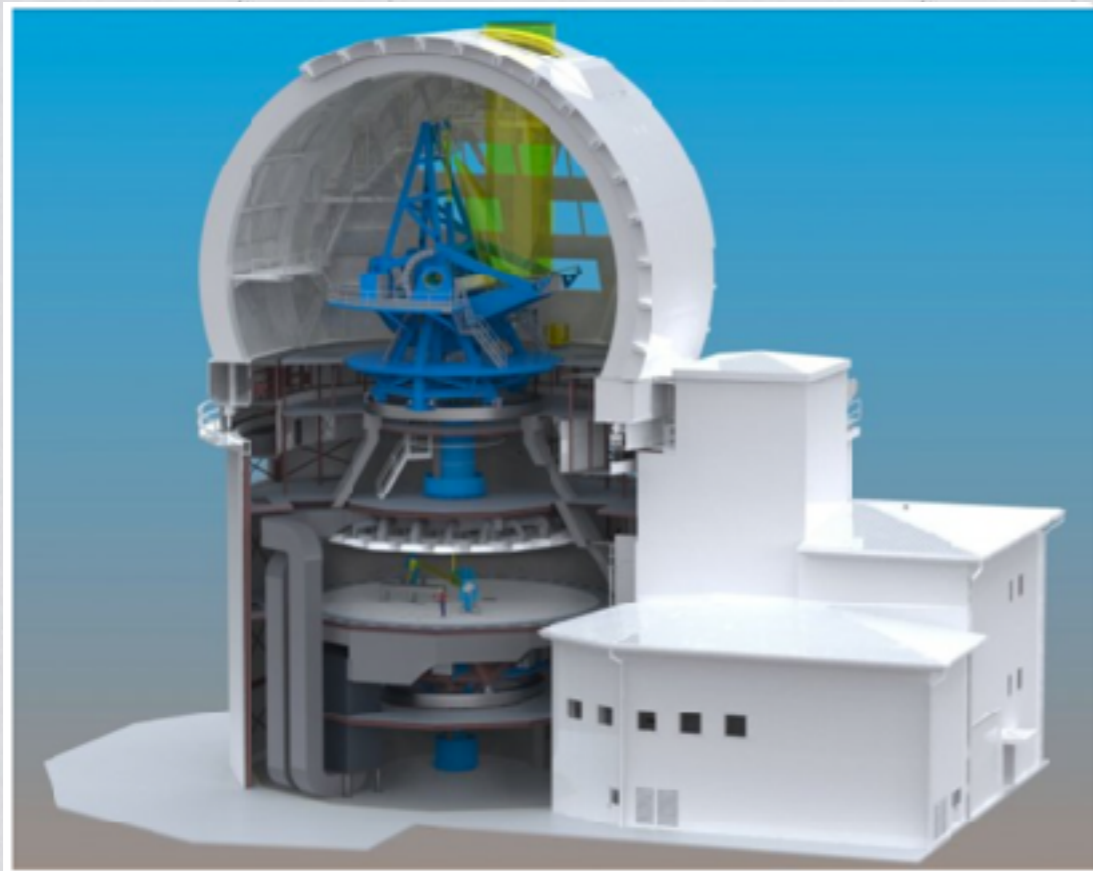


CGST 2030





# Upcoming Solar Telescopes

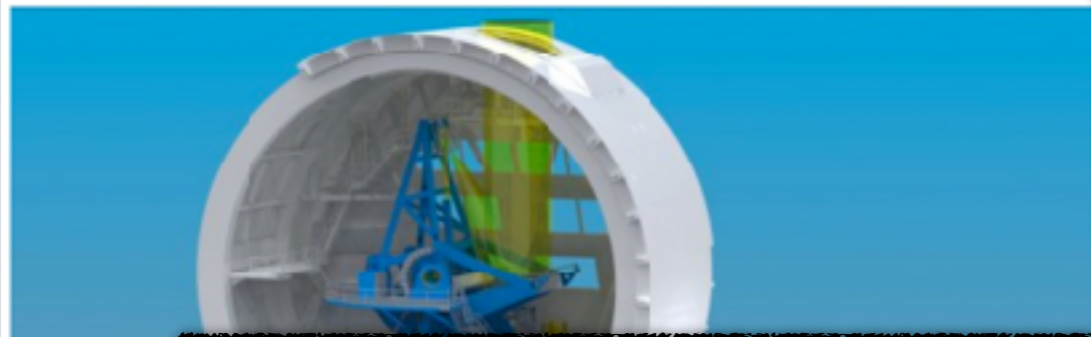


Daniel K. Inouye Solar Telescope,  
Hawaii, USA





# Upcoming Solar Telescopes



Daniel K. Inouye Solar Telescope,  
Hawaii, USA

UK and Germany are contributing to the DKIST project  
with detectors and the VTF 2D spectro-polarimeter

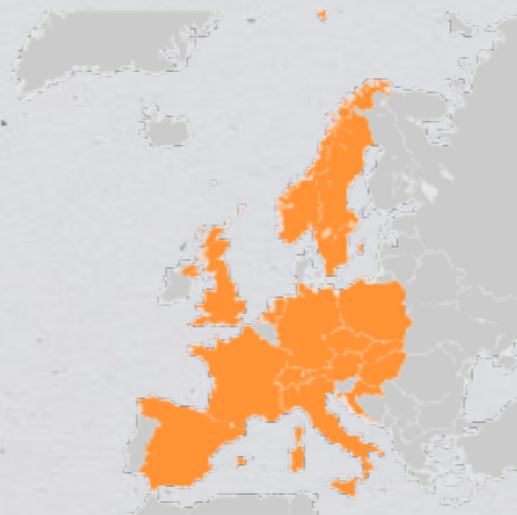















European access to observation and data





# Upcoming Solar Telescopes



Institute		Location	
<b>IGAM</b>	Institutsbereich Geophysik, Astrophysik und Meteorologie		Graz
<b>HVO</b>	Hvar Observatory		Hvar
<b>AIASCR</b>	Astronomical Institute AS CR		Ondřejov
<b>THEMIS</b>	THEMIS S.L., <sup>[note 1]</sup> INSU-CNRS, CNR		Paris
<b>KIS</b>	Kiepenheuer-Institut für Sonnenphysik		Freiburg
<b>UniDeb</b>	Heliophysical Observatory Debrecen		Debrecen
<b>INAF</b>	Istituto Nazionale di Astrofisica		Rome
<b>UU</b>	Utrecht University, Sterrekundig Instituut		Utrecht
<b>ITA</b>	Institute of Theoretical Astrophysics		Oslo
<b>IA UWrocław</b>	Astronomical Institute of the Wrocław University		Wrocław
<b>AISAS</b>	Astronomical Institute of the Slovak, Academy of Sciences		Tatranská Lomnica
<b>IAC</b>	Instituto de Astrofísica de Canarias		La Laguna
<b>SU</b>	The Institute for Solar Physics		Stockholm
<b>IRSOL</b>	Istituto Ricerche Solari		Locarno
<b>UCL-MSSL</b>	University College London - MSSL		London









# Upcoming Solar Telescopes

- \* Type: Single-sited
- \* Coordinating country: Spain

## \* Timeline

2011 - 2021  
Preparatory Phase

2021 - 2026  
Construction phase

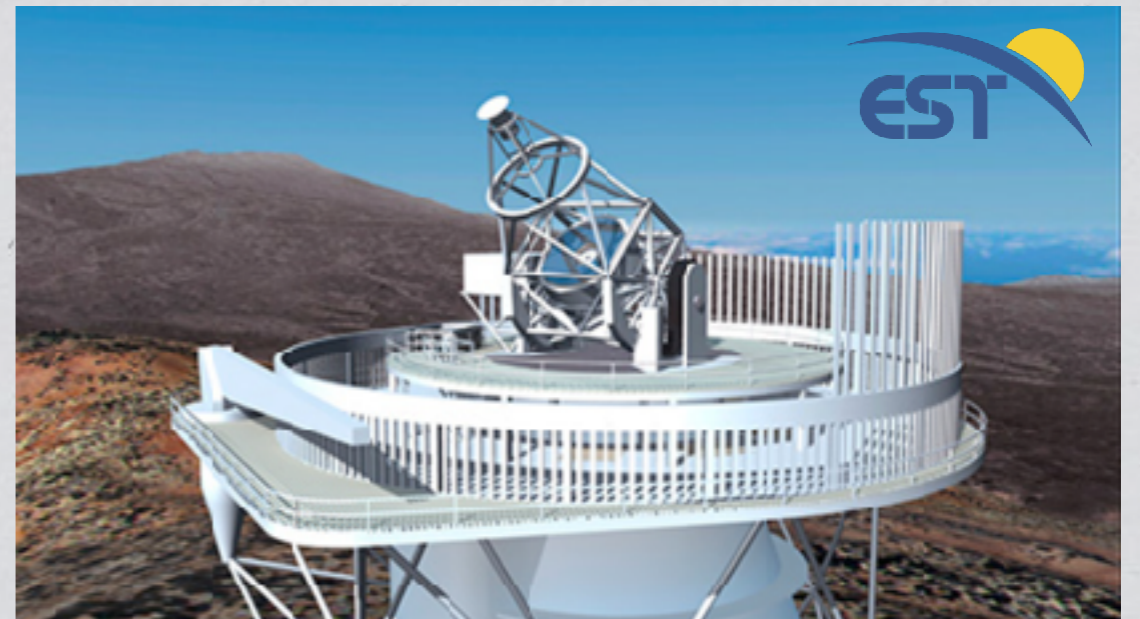
2027  
Operation start

## \* Estimated Cost

- ▶ Preparation: 10M€
- ▶ Construction: 200M€
- ▶ Operation: 9M€/year

## \* Location

- ▶ EST will be built on the Canary Islands
- ▶ Headquarters will be at the Instituto de Astrofísica de Canarias, Tenerife, Spain





# Upcoming Solar Telescopes



[...]

EST Science Data Center: EST SDC

Apart from the EST Telescope Operation and Science Centre on the Canary Islands, it is also planned to have the EST Science Data Centre in Germany, to provide data access and online services to the solar physics community. This center will provide a storage area to be accessed online through a data management system. Furthermore remote-observing facilities will be installed here.

Costs related to the installation of the telescope at the observatory form an integral part of the EST project budget. However, the construction or use of those spaces at sea-level and at mainland Europe for the EST TOSC and the EST SDC are planned to be covered by additional sources and agreements.

[...]

Online submission form: Research Infrastructure proposal to the 2016 ESFI Roadmap



# Upcoming Solar Telescopes

[...]

The EST Science Data Centre (EST SDC) will gather all expertise for producing EST science-ready data. Science-ready data will be moved or duplicated from the processing centre to the mainland Europe Virtual Observatory Compliant Data-Base (VOCDB). The SDC will be the nucleus of the scientific life of EST, where scientists are expected to come for a full data analysis and share results. If communication bandwidth allows, remote control of the infrastructure shall also be possible from the SDC. The SDC will also be in charge of the long-term data storage and the VO-diffusion of EST data. The VOCDB shall take charge of the interoperability with the VOCDB from other facilities.

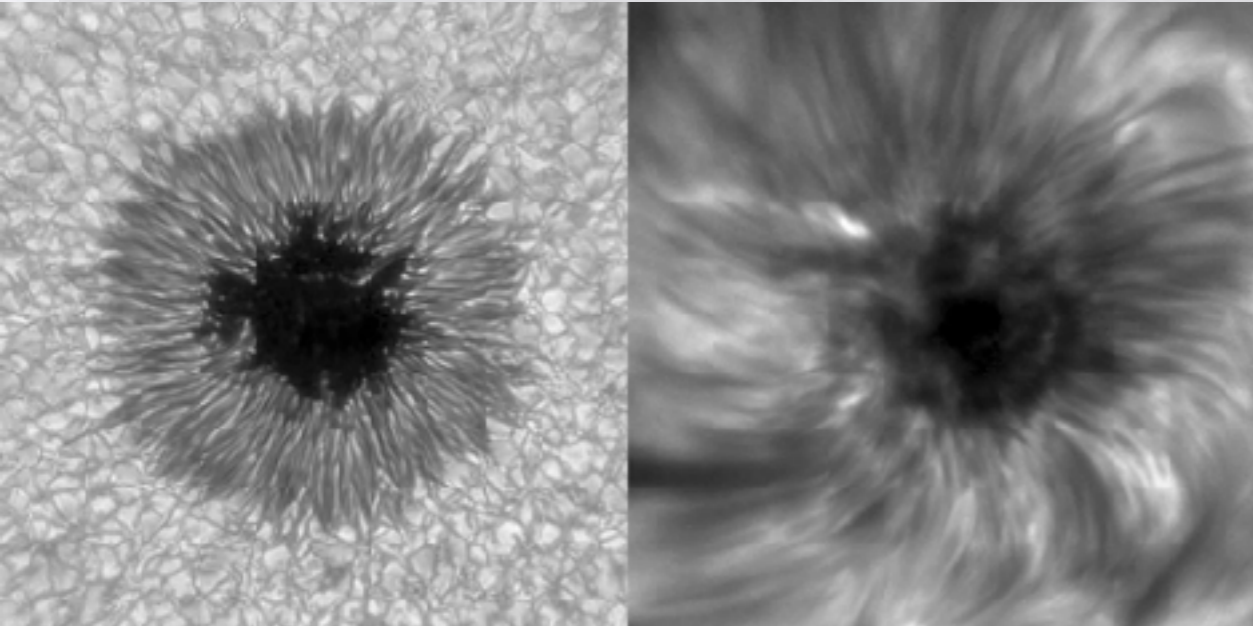
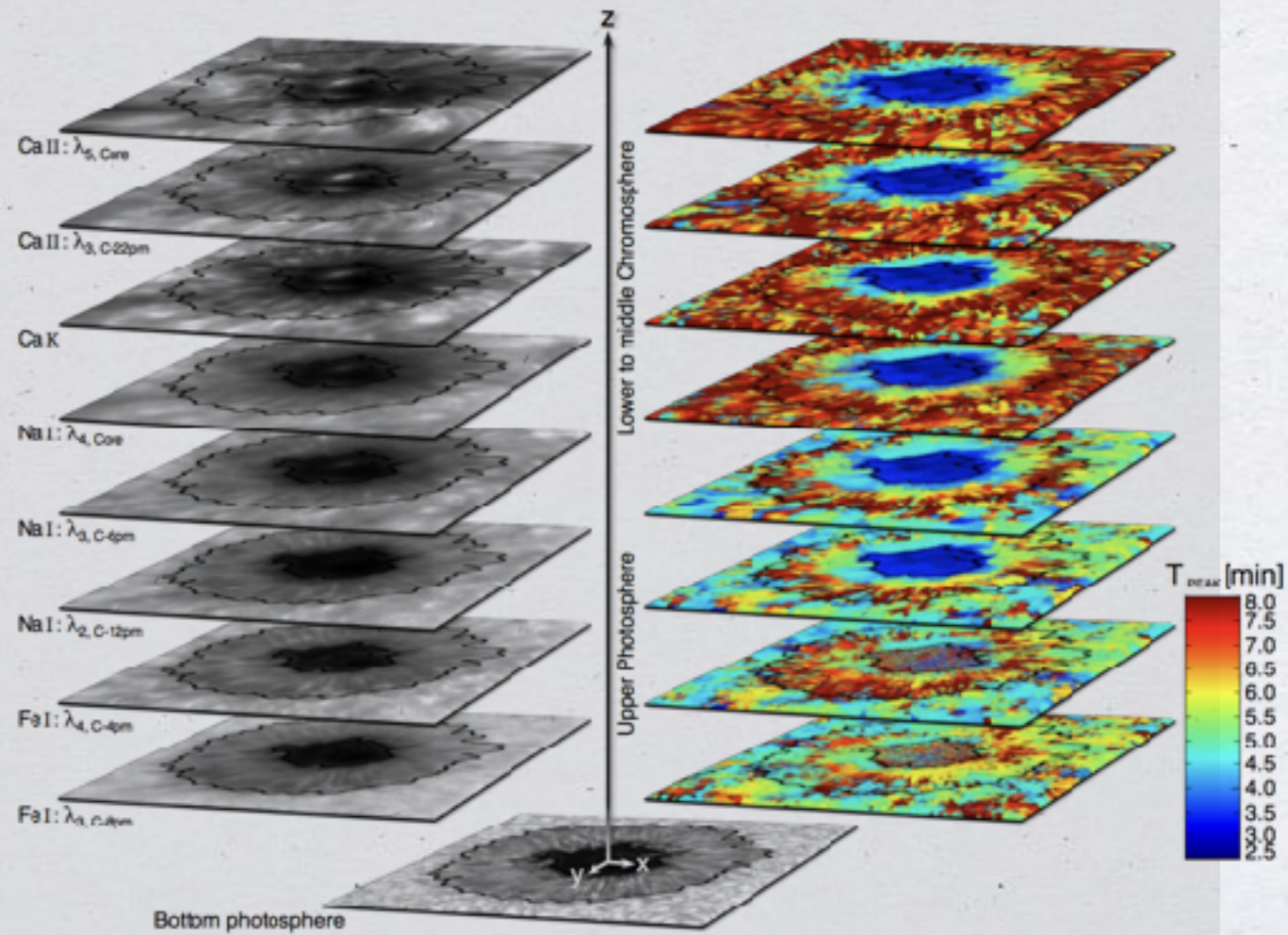
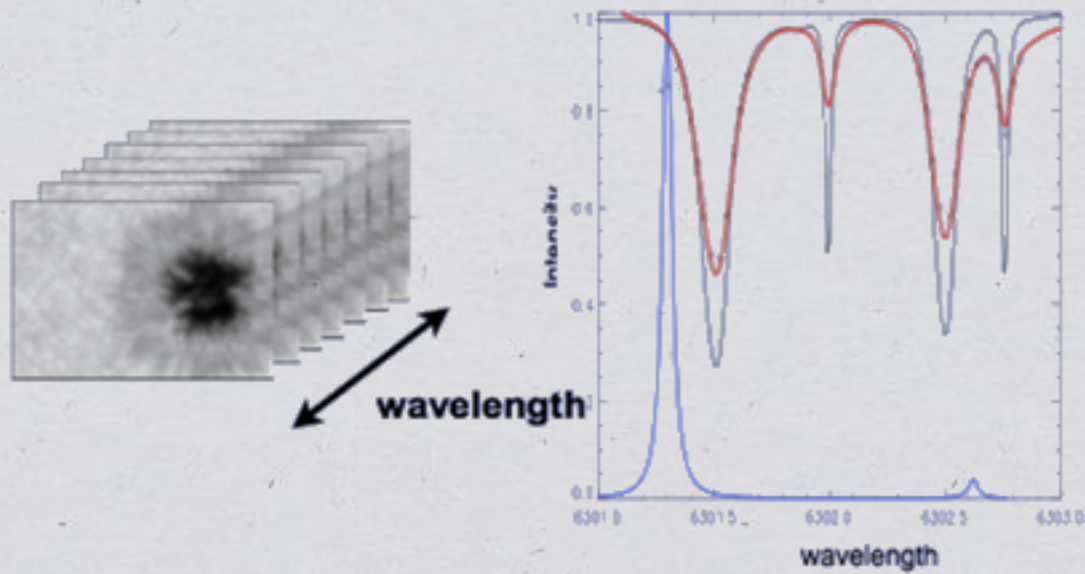
The SDC shall have offices for specialized staff in data reduction and analysis and for visiting astronomers to work on, and get familiar with, the EST data. Computing and storage capacities will be enough to guarantee the successful handling of EST data to generate innovative results.

The SDC will organize special events to gather scientific visitors there to foster discussion forums and workshops based on EST data and results.

[...]



# A Flavor of Solar Data





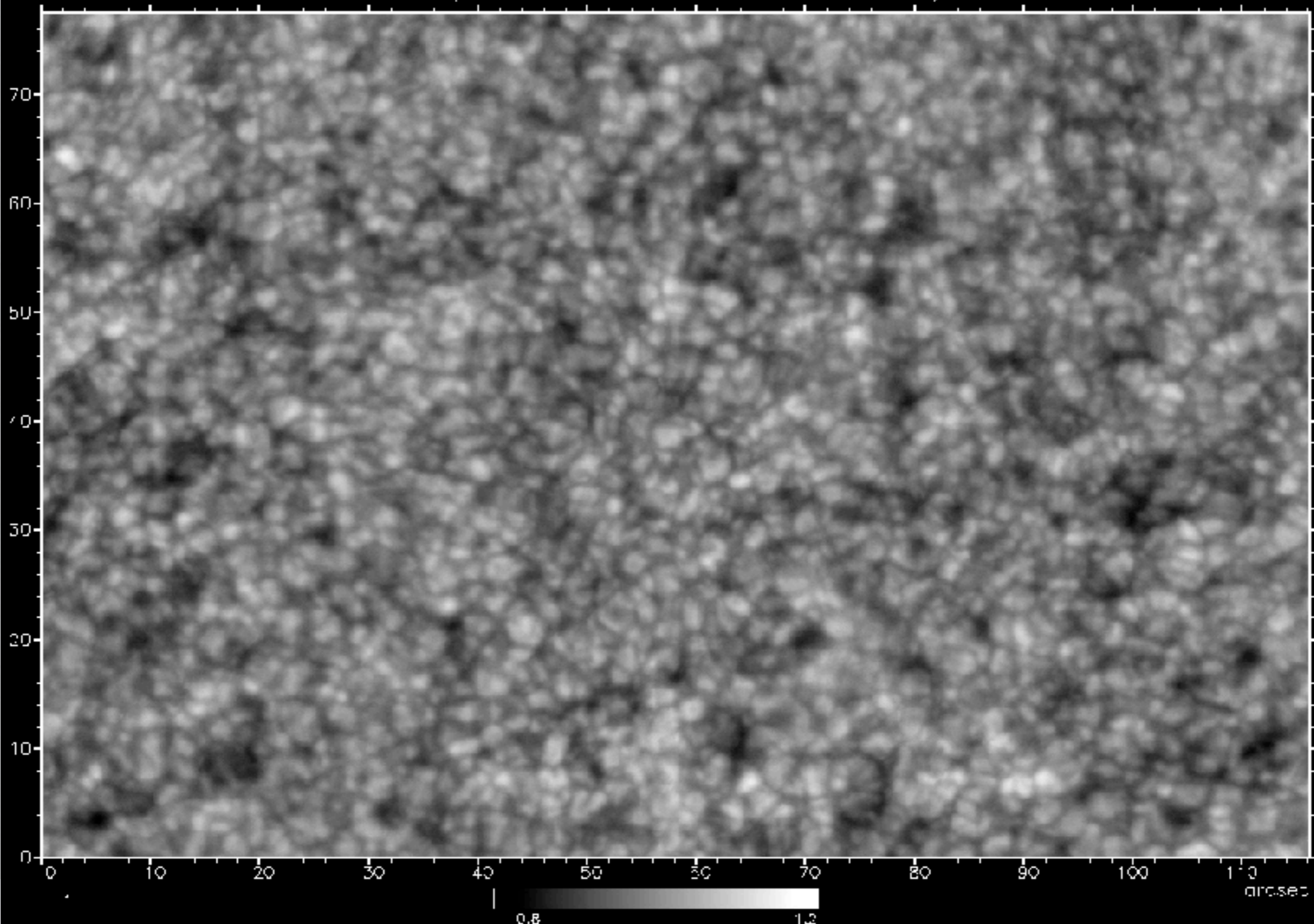
# A Flavor of Solar Data

- ◆ \* Some characteristics of ground-based solar observation:
  - ▶ Varying atmospheric observing condition (seeing).



# A Flavor of Solar Data

GREGOR B3I, center sun on 01.06.2013 08:12, 395 nm





# A Flavor of Solar Data

- \* Some characteristics of ground-based solar observation:
  - ▶ Varying atmospheric observing condition (seeing).
  - ▶ Target based (quiet Sun, sunspots, pores, plages, faculae, etc.) with a limited FOV. Pointing information become important.
  - ▶ Versatile and non-standardized observing modes as well as novel science (multi-wavelength, ...) make it difficult to unify data pipelines.
  - ▶ Upgrade might change the data characteristics for a given (upgraded) instrument.



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The major challenge for the archiving and dissemination of ground-based solar observation is the inherent heterogeneity of the data



# A Flavor of Solar Data

\* GRIS archive 2.0 (unofficial beta version)

**GREGOR Data Center**  
Search - Preview - Download

## Search GRIS

Clear all Fields!

**Observation Date (YYYY-MM-DD):**  
YYYY-MM-DD < Date > YYYY-MM-DD

**Observation Time (HH:MM):**  
--:-- < Time > --:--

**Wavelength:**  
 1083 nm  1565 nm  Other

**Observation Type:**  
 Single Map  Time Sequence

**Observation Mode:**  
 Spectroscopic  Polarimetric

**Position on Solar Disk (arcsec):**  
-1000 to 1000 arcsec < solar X > -1000 to 1000 arcsec  
-1000 to 1000 arcsec < solar Y > -1000 to 1000 arcsec

**Field of View (arcsec):**  
0 - 0 arcsec < Field of View > 0 - 0 arcsec

**Exposure Time (ms):**  
0 - 0 ms < Time > 0 - 0 ms

GDC - Gregor Data Center by [Kiepenheuer-Institut für Sonnenphysik](#). The website content is licensed.



# A Flavor of Solar Data

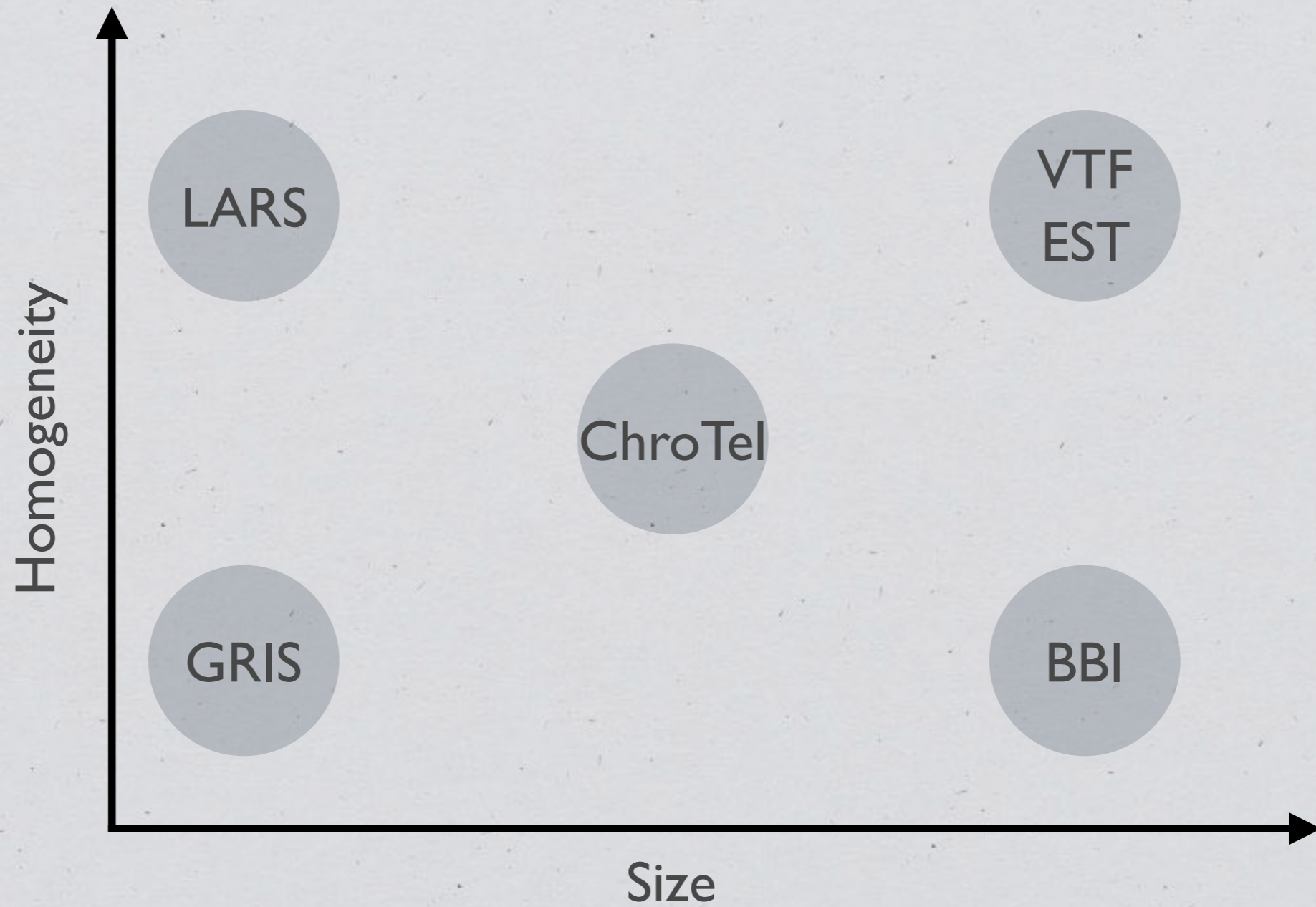
\* GRIS archive 2.0 (unofficial beta version)

**GREGOR Data Center**  
The result of your search:

No.	Date Obs	UT Start	UT End	$\lambda$ (nm)	Scan Type	Mode	Exp. Time (ms)	FOV (arcsec)	Target	Solar X (arcsec)	Solar Y (arcsec)	$\Theta$ [ $^{\circ}$ ]	Location	Map	Log File	LVL 0	LVL 1	LVL 2	Mark
1	2014-05-01	08:00:34.0	08:10:03.0	1565	Single Map	polar.	30.0	40,5	Filament/Prominence	-332	133	22,027							
2	2014-05-01	08:26:27.0	08:41:26.0	1565	Single Map	polar.	30.0	27	Filament/Prominence	3	-184	11,139							
3	2014-05-02	15:05:41.0	15:33:15.0	1565	Single Map	polar.	30.0	40,5	Filament/Prominence	-228	75	14,587							
4	2014-05-02	13:27:23.0	13:45:12.0	1565	Single Map	polar.	30.0	27	Filament/Prominence	389	-153	26,01							
5	2014-05-02	11:35:12.0	11:53:00.0	1565	Single Map	polar.	30.0	27	Filament/Prominence	96	61	6,854							
6	2014-05-02	12:07:56.0	12:26:18.0	1565	Single Map	polar.	30.0	27	Filament/Prominence	-168	151	13,71							
7	2014-05-02	14:16:47.0	14:32:14.0	1565	Single Map	polar.	30.0	40,5	Filament/Prominence	213	104	14,401							

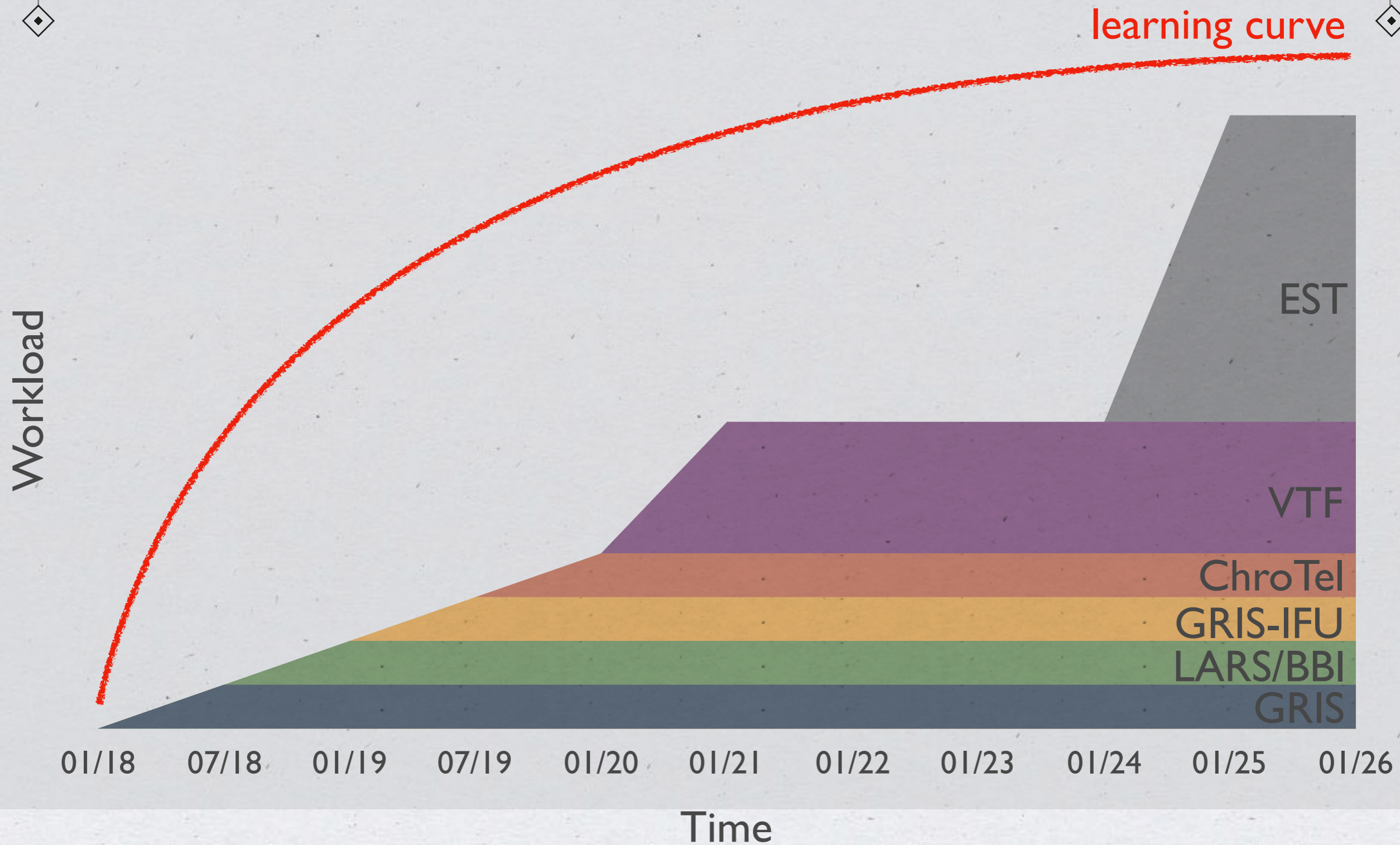


# A Flavor of Solar Data



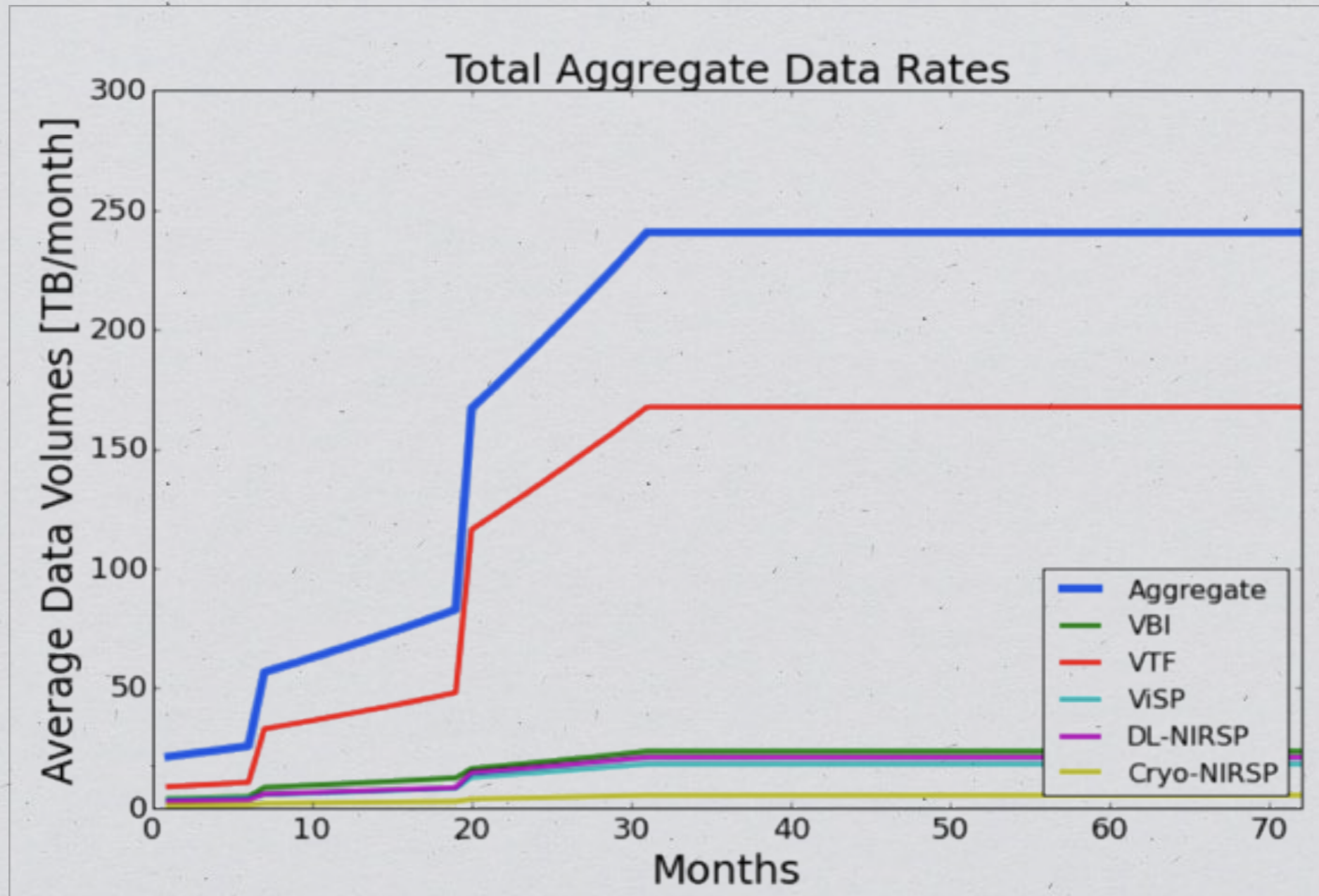


# A Flavor of Solar Data



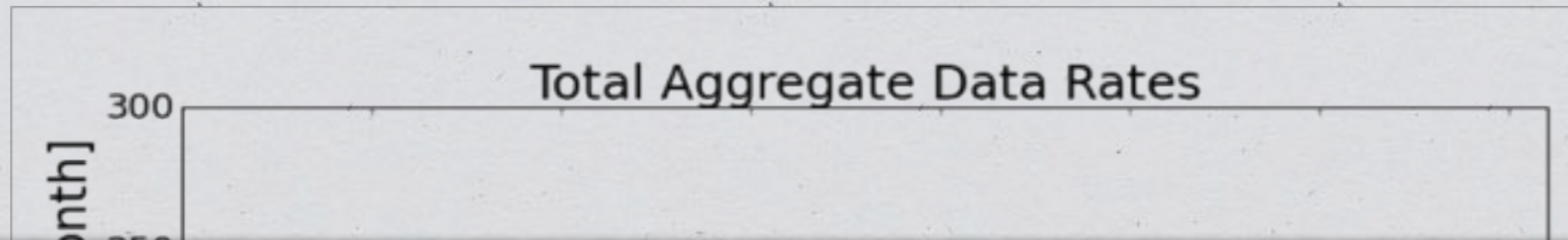


# A Flavor of Solar Data



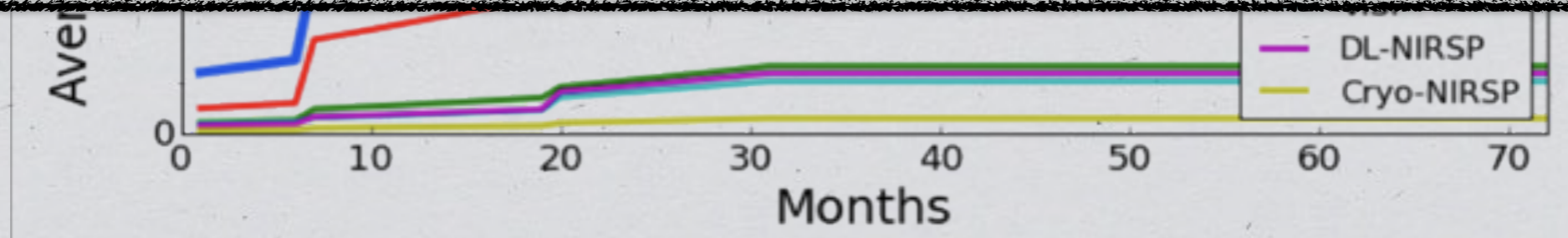


# A Flavor of Solar Data



Increase of data volume by a factor of up to 160 with respect to current ground-based observing facilities

Lessons learned with data from GREGOR and VTF/DKIST will be essential to get ready for handling EST data.





# A Flavor of Solar Data

- \* There is a need to make distribution and discovery of data as easy as possible, especially for ground based solar observation

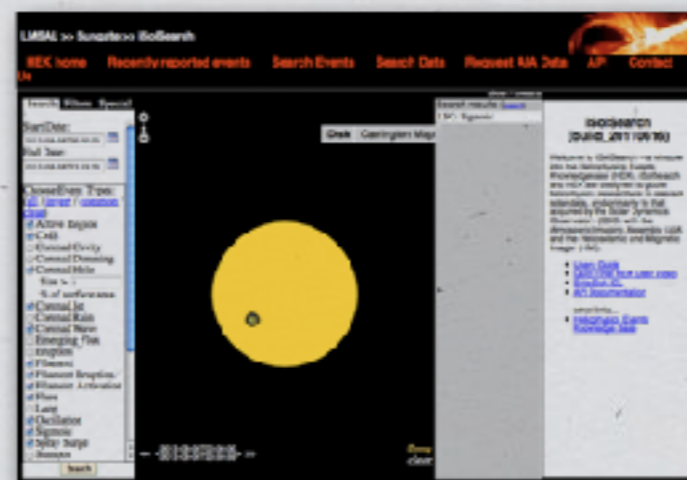
## Data archives, e.g., VSO



## Browsing tools, e.g., (j)Helioviewer

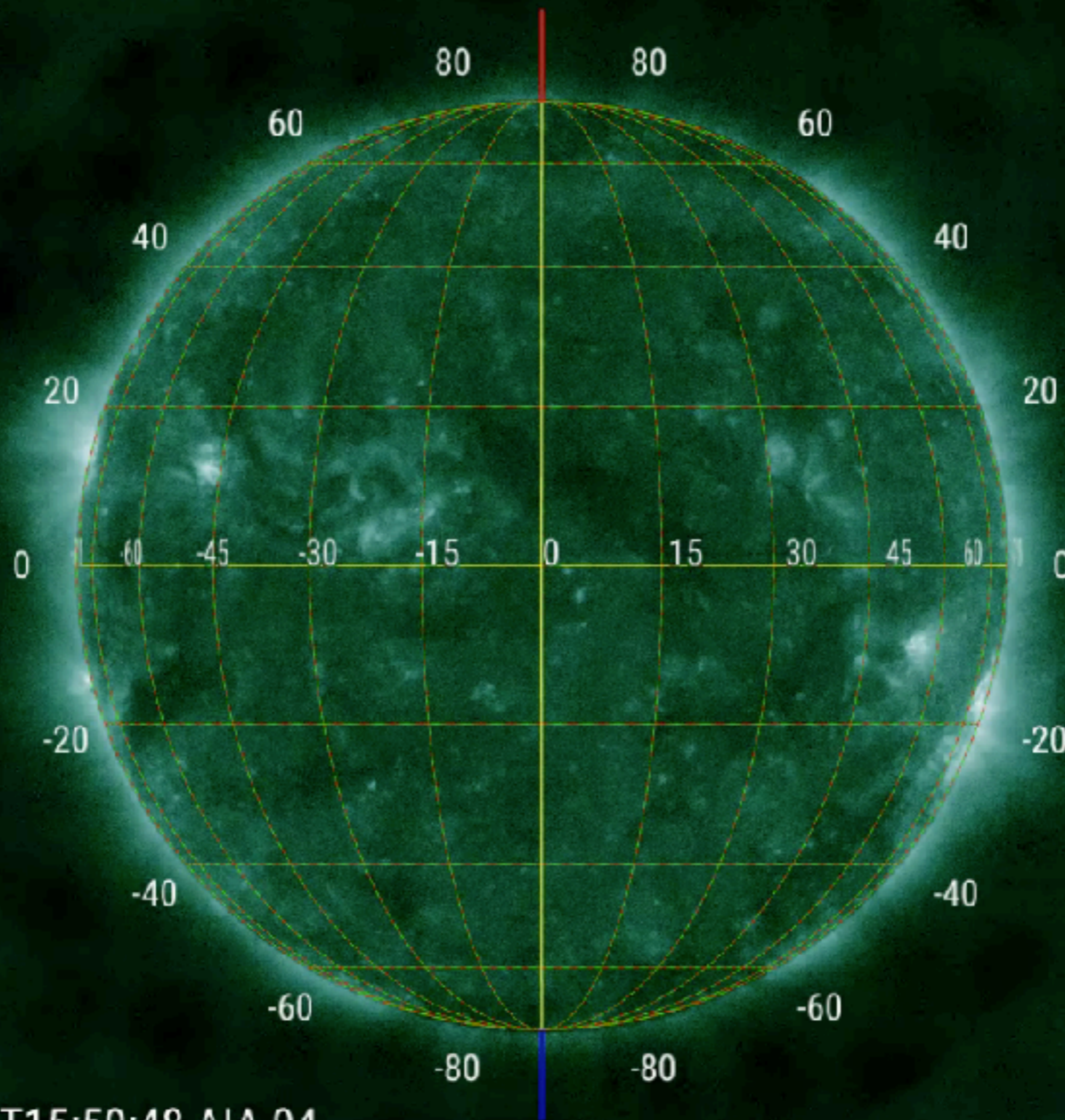


## Event searching, e.g., HEK





# A Flavor of Solar Data



2016-11-11T15:59:48 AIA 94



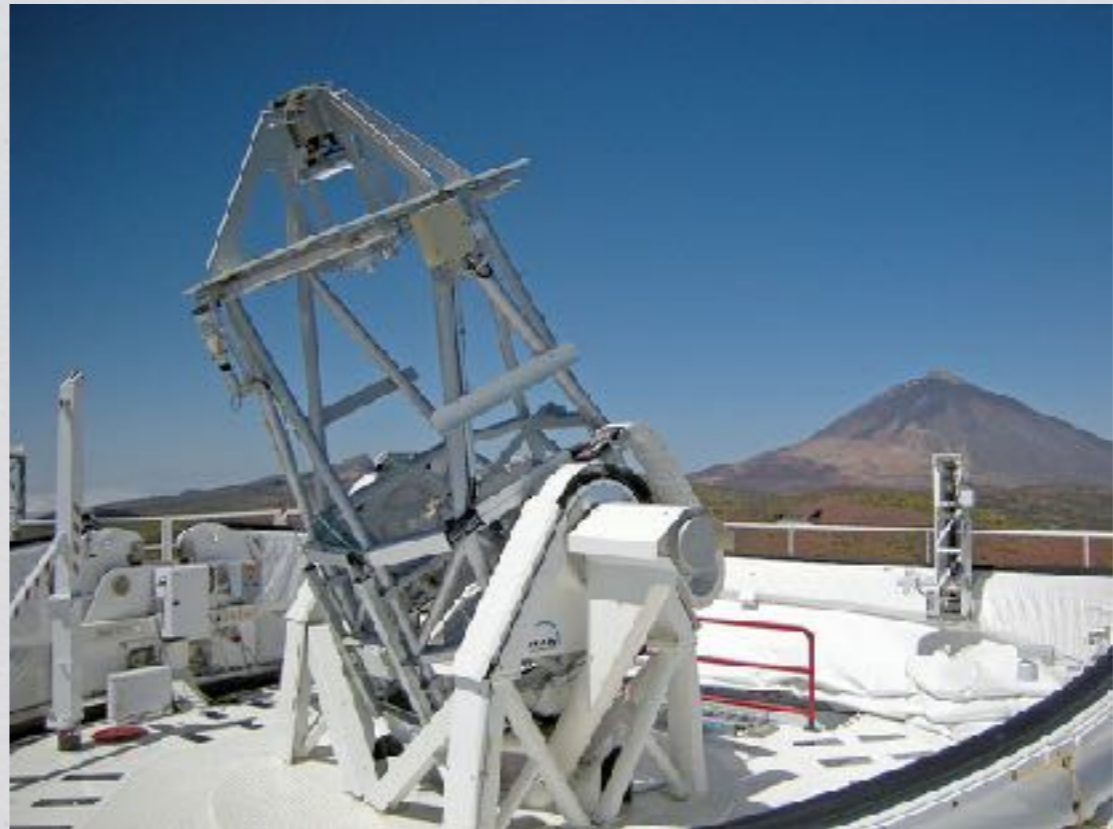
# Summary

- \* Solar physics provides input to a number of research disciplines.
- \* Space- and ground-based observatories provide complementary types of information.
- \* A new class of ground-based Solar Telescopes is on the horizon.
- \* Size and data volume of these telescopes require service mode observation and pre-defined standards for (meta)data.
- \* Challenges are the flexibility of the facilities and the subsequent heterogeneity of the data. Efforts to overcome these problems are undertaken, e.g. within the framework of the SOLARNET project.
- \* Adopting existing (meta)data standards from the astronomical community (IVOA) will be of great help for the solar community.



# Acknowledgements

- \* Thomas Hederer
- \* Manolo Collados
- \* Peter Caligari
- \* Nazaret Bello Gonzalez
- \* Carl Schaffer
- \* Philip Lindner
- \* Christian Bethge
- \* Ikrima bin Saeed
- \* Alexander Bell
- \* Andreas Lagg
- \* Svetlana Berdyugina
- \* etc.



The 1.5-meter GREGOR solar telescope was built by a German consortium under the leadership of the Kiepenheuer-Institut für Sonnenphysik in Freiburg with the Leibniz-Institut für Astrophysik Potsdam, the Institut für Astrophysik Göttingen, and the Max-Planck-Institut für Sonnensystemforschung in Göttingen as partners, and with contributions by the Instituto de Astrofísica de Canarias and the Astronomical Institute of the Academy of Sciences of the Czech Republic.